17 days until opening day

Kiss week

1



Day 16, 14.03.2019 Einstein's General Theory of Relativity, 2

hi



housekeeping

Gotta come to class

question about <u>anything</u>? I'll make a movie for you:



Madame Curie movie - we have a quorum in favor

right now: looks like Monday, March 18 Pizza FB poll established. Closes Sunday night.

Section 2 folks:

Project has begun in phases:

Document 1: software, introduction, tutorial: due March 22 Document 2: your individual dataset and project instructions: due Final Exam https://qstbb.pa.msu.edu/storage/QS&BB2019/Homework_Projects/ <u>honors_project_2019/Minervalnstructions1_2019.pages.pdf</u>

MasteringAstronomy, finally after 3 emails and phone calls:

Course ID: MABROCK41459; free code: WSSPCT-BLIDA-INANE-TOGUE-RIGOT-UNRWA check it! let me know if it is now working...

Grades to date: Projects, quizzes, notes in a pdf in the slides area the rest of your grades are in LON-CAPA or MasteringPhysics



AR and MRS MINIVER t

March 2019







"'rinciple of equivalence"



6

Equivalence Principle

his first aha moment

there's something the same about acceleration and gravitation

stupid elevator trick, #1 gravitational attraction

stupid elevator trick, #2 gravitational attraction



identical



The Equivalence Principle demands that

If some phenomenon happens in a gravitational field,

then it must happen an accelerated frame.

and

If some phenomenon happens in an accelerated frame field, then it must happen a gravitational field.

acceleration warps space

from the Equivalence Principle



gravity should warp space

9

a beam of light in a high gravitational field

should appear to bend

acceleration warps time

from the Equivalence Principle



gravity should warp time

a clock in a high gravitational field

should appear to run slower than a clock in a weaker gravitational field

"gravitational red shift"

gravitating bodies...masses:

warp both space and time.

They warp: spacetime

furthermore!

indistinguishable

free fall

in a gravitational field eliminates gravity







empty outer space without any gravitational fields

the free-fall recognition became:

Maybe gravity is not a force at all?

modeling this was heroic

the most technically challenging piece of physics ever

Hilly 105 $\frac{d}{dt} \cdot 2\dot{f} = -\int \sigma \frac{\dot{s}}{2\pi} = -\int \sigma \frac{d}{dt} \left(\frac{t}{2}\right)$ $2\dot{f} = \int \sigma \cdot \frac{t}{2\pi} + C$ Fin Die Bewegungs gleichungen no moterillen Tweedles landen : ales (bei west): - d & 3 + 2, + . + 9 = - 1 > 2 = 2 a. thorne 1/2 12 and Dos Reptorat ! E= c, (1- 2 2+ 2 2) -12 4= 22 -=1-====+==== , An herys ige time : forman = + m (qu, dx. + ... + quy dx v) also E = - gr dt 2 Ester Angl A (T + x2 1) Perikelburyung bei ally. on pro po : C.E . 20:-0: 4 + 2" 2'd q + da' = 2 c. E - 2 c. + C. A Datie ist (and flow S. 6 +, and S.7 gl. 5) , w 2"dy 2: (Jo : de + C)'de du Oring I und m 2 it ald mendlich Alin er · / 9. x + . + . + . $\frac{d^{2}T}{dx^{2}} = \frac{1}{r}\frac{dT}{dr} + x^{2} \cdot \frac{1}{r}\frac{d}{dr} \left(\frac{1}{r}\frac{dT}{dr}\right)$ 9 ... Bw 2(1.21, 12)2"dy" = (fo 1 + C)2(22dy + dz") - guy dt = - gry /1 $d = \left[\frac{2c_0(\delta - c_0)}{c_0 A_1} \right] 2^* dy^* = \left(\int_0^{\infty} \frac{1}{2} dy^* = \left(\int_0^{\infty} \frac{1}{2} + C \right)^2 dx^*$ $\Delta T = \frac{3}{r} \frac{dT}{dr} + \frac{1}{r} \frac{d}{dr} \left(\frac{1}{r} \frac{dT}{dr} \right)$ with buincheichligt Dass sich am Folde unt Du East wich Eyz' = 39. $\frac{2}{2x}\left(x^{2}\frac{N}{r^{2}}\right) = 2x\frac{N}{\mu^{2}} + x^{3} \cdot \frac{1}{r}\frac{d}{dr}\left(\frac{N}{r^{2}}\right)$ 2 x dr (+ the Cortable with Jeso Wer : O ge sign sind , were auch wide for x, = x , x = y $d p = \frac{\int_{-\pi} \frac{1}{2} \cdot \mathcal{C}}{\sqrt{c_0 (z_1 (b \tau_0) + \frac{1}{2} t_0^2) \cdot z^4 - (S_0 + \mathcal{C}_2)^2}} dz$ W . 41 -9' + C $\frac{2^{2}}{2x^{2}} = 2 \frac{W}{r^{2}} + 5 x^{2} \cdot \frac{1}{r} \frac{d}{dr} \left(\frac{W}{r^{2}} \right) + x^{4} \cdot \frac{1}{r} \frac{d}{dr} \left(\frac{1}{r} \frac{d}{dr} \left(\frac{N}{r^{2}} \right) \right)$ ds = Vg, x + g ... + 2g, x + ... 2g, x + ... + g. " " " B' (c'-q'- c' + n c' d') (6 4 1- d. c. (1- d + n d') = c. (1-2 d 2 = x2y. 1 d (N/r2) Die group and aber , his and weadle the torede Ording $q^{2} = c_{0}^{*} \left(1 - \frac{d}{2} + h \frac{d}{2^{*}}\right) - \frac{c_{0}}{f^{*}} \left(1 - 2 \frac{d}{2} + \frac{f^{2}}{2}\right)$ gry = Co (1 - 2 + 3 12) Dola whereint and De garage a = $x^2 \cdot \frac{1}{r} \frac{d}{dr} \left(\frac{M}{r^2} \right) + x^2 y^2 \cdot \frac{1}{r} \frac{d}{dr} \left(\frac{1}{r} \frac{d}{dr} \frac{g(M)}{r^2} \right)$ $\frac{dt^{*} + t^{*} dy^{*}}{dt^{*}} = C_{o}^{*} \left\{ 1 - \frac{C_{o}^{*}}{C_{o}} + \frac{d}{2} \left(2 \frac{C_{o}^{*}}{L^{*}} - 1 \right) + \frac{d^{*}}{2^{*}} \left[n - \frac{C_{o}}{C_{o}} \right] \right\}$ $\Delta(x^{2}\frac{N}{r^{2}}) = 2\frac{N}{r^{2}} + \int_{0}^{\infty} x^{2}\frac{1}{r}\frac{d}{dr}\binom{N}{r^{2}} + x^{2}\frac{1}{r}\frac{d}{dr}\binom{1}{r}\frac{d}{dr}\binom{1}{r}\frac{d}{dr}\binom{1}{r}$ $\frac{d}{dt} \left[\frac{d}{dt} \int_{0}^{t} \int_{0}^{t} \frac{d}{dt} \int_{0}^{t} \frac{d}{dt}$ $\frac{3}{r}\frac{dT}{dr} + \frac{1}{r}\frac{d}{r}\left(\frac{i}{r}\frac{dT}{dr}\right) + 2\frac{N}{r^2} + x^2 \int_{0}^{\infty} \frac{1}{r}\frac{d}{dr}\left(\frac{M}{r}\right)$ $\int_{1}^{1} \frac{dt (1 + x' \frac{1}{2})}{\sqrt{(2-2)(2-2)(2-2')(2-2')}}$ $\frac{2\pi}{V_{2,1_{1}}} \left\{ 1 + \frac{1}{2} \left(\frac{1}{2} + \frac{1}{2} \right) \left(x + \frac{t' + 2}{2} \right) \right\}$ H $\int_{-\infty}^{\infty} \frac{d^2}{2} \left[dz^2 + z^2 dy^2 \right] \left(1 - 2\frac{d}{2} - \rho \omega y \frac{d^2}{2z} \right) = \int_{0}^{\infty} z^2 dy^2 - 2 - 2 + \frac{1}{2}$ 2'+2" = - J- & / C + = - J. -#2 -+ # # 2 d2' (() (1-2' - popl') = dy = jor's [. $+2, +2, +2'+2'' = \frac{A}{2z(c_{*}-b)}, 2, +2, \frac{A}{2z(c_{*}-b)} + \frac{A}{C}$ + 1/2 (1 1/4 - 1/6 y de (x/da) - x de (y/da) = d dy x-xy = de 60 6 (1-2 + 120) 1,2, = (1+the the 1- 2 0 2"+ A (21+1) 4 2 3- [A" (n+1) $\alpha \varphi = \frac{\mathcal{B}}{\delta l \overline{\epsilon}} \int \frac{1}{\sqrt{-z^2 + (\frac{1}{\epsilon} + z)} d z^2 - \frac{1}{\epsilon} \tilde{\omega} \frac{\partial}{\delta}}$ (ENCO) dr()====== 6 ds - 3 + 1 + 1 + 1 - 5 + 7/ 1=-14 died abed De grossen Vingla > 2, 12 plats , Si bester Klan 2 mil 2' -1+L 1 = 3 4 Inhormant mon 2'+2 2 A , 2,+22 = [(+ OA) = ((1-2) ÷ (1+ 1) $2_{1} 2_{1} = \frac{1}{2} \left[h(\frac{1}{2}) + (n+1) A \right] - \frac{1}{2} A = \frac{1}{2} \left[\left(\frac{2}{2} \right) c + (n-1) A \right]$ $\frac{1}{2} \int \frac{1}{2\pi} \frac{1}{2\pi} \int \frac$ + (+ # : + 2) (+1

there should be observable consequences to the Equivalence Principle

and Einstein knew it

and calculated them - half a decade of Newton-like concentration

Einstein had to learn that geometry & <u>energy-mass</u>

interact & that space and time respond

After his "happy thought": 5 years for him to figure it out



he had to go back to school...privately with his buddy Marcel Grossman





tests of general relativity

There are a handful of "classic tests"

of these ideas:

that space and time are warped by gravitation



light beam what about time?



В

use a clock Gravitational Red Shift is built so B moving away from A A and B under Orevitour phone ScGaPSger inertial frames at each ifiyou get where you want to go you just confirmed General Relativity per second

force up at g

receives at say 5 ticks per second

1 second

"Advance of the Perihelion of the Orbit of Mercury"

Vulcan?

Mercury

misbehaves

"advance of the perihelion"

Einstein calculated it including the sun's warping of space



1916: Got precisely the right amount.

Heart palpitations when he scribbled the result on his paper...

point of closest approach of the orbit advances by 43 seconds of arc per century

the mother of all experiments

the "solar eclipse" experiment

Google solar eclips All Vide

About 6,480,0

Videos



Solar eclip experiment

KSDK News YouTube - Aug

People al

What is th What fam Who prov How can

A Picture-

Nature: Total

Solar Eclipse Model I Science project I Education.com https://www.education.com > Entire library > Science projects > First Grade > Science * Check out this fun science fair project idea for 1st grade: a solar eclipse model that demonstrates total, ... In this experiment, you'll use an apple as the sun.

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-Perfect Sola	^r Eclipse Experime	ent - Sky & Te	lescope	

https://www.skyandtelescope.com/...news/a-picture-perfect-solar-eclipse-experiment/ -Feb 16, 2018 - In the August 2016 issue of Sky & Telescope, I explained that I was going to set up an experiment to measure the gravitational deflection of ...

How a Total Solar Eclipse Helped Prove Einstein Right About ... https://www.space.com/37018-solar-eclipse-proved-einstein-relativity-right.html 🔻

May 29, 2017 - The event will be a great opportunity to revisit a groundbreaking experiment that occurred during a total solar eclipse, and helped confirm Albert ...

Solar eclipse of May 29, 1919 - Wikipedia

https://en.wikipedia.org/wiki/Solar_eclipse_of_May_29,_1919 -

A total solar eclipse occurred on May 29, 1919. With the duration of totality at maximum eclipse of 6 minutes 51 seconds, it was the longest solar eclipse since ...

Greatest eclipse: 13:08:55 Duration: 411 sec (6 m 51 s) Max. width of band: 244 km (152 mi)

The Solar Eclipse and Eddington's Experiment - Department of Physics

https://www.cmu.edu/physics/news-events/news-archive/.../eclipse-eddington.html -Aug 15, 2017 - At 1:45 on Monday afternoon, August 21, the moon will begin blocking the light coming to us from the Sun. Over the next half hour, the skies will ...

"Solar Eclipse Experiment"



light

obeys the strong Equivalence Principle

the laser pointer...for real

The star is actually at A

Zariel. 14. X. 13. Each quelester Hers Kollege! time surfache theoretische Ufer legung market die Annahmes plausitel, dass Lichtstrahlen in einem Geavitation felde eine Deviation uphren. 15 Lechrebahl An Somencande misste diese Ablenkung (R Stationada and mil 1 abuchunen 50.84 To ware deshall von grösstem Interesse, bis que une grosses Sommen whe fills Firsteine bei turendung der stänkesten Vergrösserungere bei Tage (ohne Somenfinsternis) gerehen werden komun

1911 calculation – initially wrong, only the E=mc² component...

In 1915 he changed his 1911 calculation to include the warping of space...worth x2

The deflection should be about 1/4 milli-degree



eclipse experiment May 29,1919

Sir Arthur Eddington led 2 teams:

Gulf of Guinea

& Brazil





there was some cloudiness!

Eddington had 10 seconds to get a photograph







1/16 plates had usable data

Eddington announced the result

November 6, 1919 at the Royal Astronomical Society meeting

Einstein woke up in Berlin the next morning and was famous.

eclipse announcement at scientific meeting, 11/06/19:

instant celebrity, 11/07/19

cover of December 14, 1919 issue of *Berliner Illustrirte*



caption: "A new great figure in world history: Albert Einstein, whose investigations signify a complete revision of our concepts of Nature, and are on a par with the insights of a Copernicus, a Kepler, and a Newton."



the Times golf editor

New York Times, November 10, 1919

One of the speakers at the Royal Society's meeting suggested that Euclid was knocked out. Schoolboys should not rejoice prematurely, for it is pointed out that Euclid laid down the axiom that parallel straight lines, if produced ever so far, would not meet He said nothing about light lines.

Some cynics subgest that the Einstein theory is only a scientific version of the well-known phenomenon that a coin in a basin of water is not on the spot where it seems to be and ask what is new in the refraction of light.

Albert Einstein is a Swiss citizen, about 50 years of age. After occupying a position as Professor of Mathematical Physics at the Zurich Polytechnic School and afterward at Prague University, he was elected a member of Emperor William's Scientific Academy in Berlin at the outbreak of the war. Dr. Einstein protested against the German professors' manifesto approving of Germany's participation in the war, and at its conclusion he welcomed the revolution. He has been living in Berlin for about six years.

When he offered his last important work to the publishers he warned them there were not more than twelve persons in the whole world who would understand it, but the publishers took the

New York Times, December 3, 1919



spondent called at his home to gather from his own lips an interpretation of what to laymen must appear the book with the seven seals. Dr. Einstein him-



now recovered from exhaustion and photogenic: 1920

Gravitational Lensing - an off-hand prediction of Einstein

Foreground objects can distort, and magnify distant background galaxies.

SDSS J1420+6019

SDSS J1627-0053

gravitational lensing now a tool

no longer a discovery!

Second view of this:

"The Hafele-Keating experiment"

an atomic clock was carefully carried around the world in 1972 and carefully calibrated and compared with groundbased clocks

correct

Predicted Effect	Flying East	Flying West	
GTR (Gravitation) STR (Velocity) Total	+ 144 ± 14 ns - 184 ± 18 ns - 40 ± 23 ns	+ 179 ± 18 ns + 96 ± 18 ns + 275 ± 21 ns	
measured:	- 59 ± 10 ns	+273 ± 7 ns	

redone twice more in airplanes and rockets/satellites

J. Hafele and R. Keating

about half of their effect was due to the gravitational difference between Earth and the flight's altitude

spacetime in general relativity

Einstein got rid of gravitational forces in GR

Masses warp spacetime...

Since the shortest distance between two spacetime points is a light-path, this "maps" the shape

In GR gravity is not a force, but a "topography" of spacetime that forces objects to take the shortest curved path in spacetime

Earth's orbit is then just us following the shortest distance around the sun...not a gravitational force

okay.

Spacetime might be curvy, bumpy, ..."warped"

a "non-Euclidean geometry"?

Euclid's Geometry starts with 4 terms and 4+1 postulates:

Point, Line, Plane, Space

- A straight line can be drawn between any two points 1.
- 2. A finite line can be extended infinitely in both directions
- A circle can be drawn with any center and any radius 3.
- All right angles are equal to each other 4.
- Given a line and a point not on the line, only one line 5. can be drawn through the point parallel to the line

a System of a series of proofs, each building on the previous, to a whole system of mathematics

like, actually... a lot of algebra problems before algebra was invented
Einstein's mathematics of GR

led him to have to consider non-Euclidean Geometries which were still timidly being studied by mathematicians

not so far-fetched

we live in such a geometry



what's a ''straight line''



on a sphere?

shortest distance between 2 points

http://gc.kls2.com/



Euclid's 5th Postulate

parallel lines never meeting?

only in a flat space

sum of interior
angles in a triangle
= 180°?

only in flat space

on sphere > 180°



"warping"

means that geometry

spacetime geometry

mixes with mass, energy, and pressure

General Relativity

Einstein's GR equation

complicated mathematics geometry of spacetime

 $R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = \frac{8\pi}{c^4} T_{\mu\nu}$

we'll call it: "G = T "



 \leftrightarrow

mass-energy, pressure, & momentum



Einstein grossly underestimated

the richness of his theory

he knew he'd exhausted the possible solutions to the GR equations

He was wrong...and irritable about it

wrong. Almost immediately:

from the foxhole, 1915



Karl Schwarzschild, 1873-1916

Yes. I mean *from* a foxhole.

The **first exact solution** to GR...Einstein had used some approximations for light-bending, etc.

The equations of spacetime outside of a spherical mass.

a big mass.



escape

Suppose a rocket is shot straight up... when it goes "ballistic" (no propulsion)...what happens?

It depends.



More initial velocity, the more likelihood that the rocket will escape the pull of the Earth's gravity.

This happens when the kinetic energy = potential energy

$$v_{\rm esc} = \sqrt{\frac{2GM_E}{R_E}}$$

From Earth: 11.2 kilometers per second...~25,000 mph

 $v \approx v_{esc}$

what about light?

suppose the question is not:

"What's the escape velocity from a sphere of mass M?"

BUT

"What's the radius of a mass M for which the escape velocity is = *c*?"

 $v_{\rm esc} = \sqrt{\frac{2GM_E}{R_E}} \quad \longrightarrow \quad c = \sqrt{\frac{2GM}{R_S}}$

R_S called the Schwarzschild Radius

$$R_S = \frac{2GM}{c^2}$$

It seemed to be a magic radius...



46

the

Schwarzschild Radius falls out of his solution to G.R.

it's not likely

R_s is incredibly small

and density, incredibly high All of the mass of:



Impossible, right?...thought to be a disaster for the theory.

inside of R_S :





everyone fretted over this for more than a decade

1932, Georges Lemaître found that a slight change of coordinate axes

changed the problem completely

Black Holes

The Schwarzchild Radius was not a flaw in the theory

simply an insult from Nature!



Einstein calculated that the normal formation of a star of gravitational accretion could never form in so small a volume...and stars get bigger not smaller, right?

1939: Robert Oppenheimer & Hartland Snyder showed how.

5¢ worth of stellar physics no charge

Hertzsprung-Russell Diagram...aka H-R Diagram



5¢ worth of stellar physics no charge

Hertzsprung-Russell Diagram...aka H-R Diagram



stars radiate energy – that's their job!

being stable is their challenge...



The "main" sequence

a balancing act

inward pressure from gravity

VS

outward pressure from radiation



A star's fate is determined by how massive it is.

gravity pulls core/atmosphere: in Radiation pressure from nuclear fusion in core: out

H begins to "burn" to He



pp cycle







Helium-capture reactions



Other reactions



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a balancing act

inward pressure from gravity

VS

outward pressure from radiation





WINS

STOPS



and then a special effect takes over:

It stops abruptly in seconds

Explosively.

gravity pulls core/atmosphere: in

Radiation pressure from nuclear fusion in core: out

$e + p \rightarrow n + \nu_e$ everywhere...

the star shrinks dramatically

neutrons cannot all be on top of one-another



supernova!









SN 1993J M81



Crab Nebula...supernova remnant from 1054 AD

first infant picture of SN

20 September 2016

Victor Buso:





NGC 613 65M ly away

aftermath of a SN

mass-ejection and a neutron star



"pulsar"...a rapidly rotating neutron star: few milliseconds to seconds in rotation rate

The source of all elements < Fe. We are made of star-stuff

Cluster Lensing And Supernova survey (CLASH) SN Candidates in 25 Clusters (Of the 30, ~30% are Type Ia) **SN "Galb** l "Ántonius Pius"

SN "Augustus"







"Claudius"

Discover

SN "Titus'

Reference







"Hadrian'







"Marcus Aurelius



"Crimson

"Burgundy

what if $M > 3-15 \times M_{sun}$?

Nature turns viscous



very peculiar

Gravity wins. Nothing gets out, not even light:

BLACK

no light

the most extreme warping of spacetime in Nature



outside of $~~3R_{\rm S}$

a black hole behaves like a normal object with Newtonianlike gravity

So, how are they found?

Because they're hungry.



the matter sucked in accelerates... and accelerating charges do what?

Radiate...**X-Ray, radio frequencies** typically

Three kinds:

- Stellar black holes 100's found with Hubble 1.
- **Supermassive black holes** seems that all galaxies 2. have one: billion's of stars' worth
- **miniature black holes**. complete speculation, a gleam 3. in some theorists' eyes

Stellar Black Holes

many stars' fate

black hole anatomy



simplest black hole: not rotating and not charged

realistic black hole: rotating...Kerr Black Holes



"singularity"

"photon sphere" $1.5 \times R_{\rm S}$

67

often in binary systems with characteristic X-ray emission Cygnus X-1/HDE 226868 1972





what if you fell in?

everyone asks this!

from the outside:

from the inside:



also, atomic spaghetti

striking: time slows near the ultra strong gravity

69

black holes are in the news and on the silver screen



Natch in: 🗮



Embed Code

Supermassive

Black Holes center of every (?) galaxy





100,000ly-ish
let's orient ourselves

Sagittarius A and A*





https://www.reddit.com/r/eliteexplorers/comments/92li23/galactic_map_with_labeled_regions/ 73

nunity expedition:

Milky Way the in

news

enormous, central black hole confirmed



S2:

- has a period of 15 years and comes close 17 light-hours so center object is smaller than that
- orbit implies M(center) = 4.1M solar masses



S14:

• comes closer - 6.25 light-hours - so center object is smaller than that!

Galactic black holes:

Milky Way

 $4 \times 10^6 \times M_{sun}$

M84

 $300 \times 10^6 \times M_{sun}$

M87

3.5 x 10⁹ x M_{sun}



Hubble Space Telescope PRC97-12 • ST Scl OPO • May 12, 1997 • B. Woodgate (GSFC), G. Bower (NOAO) and NASA

M87 Active Glactic Nucleus (AGN)



